

Designing a radiative antidote to CO₂

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Key Points:

- Spectrally-flat solar geoengineering reduces CO₂-induced change in mean rainfall disproportionately more than mean temperature.
- A spectrally-tuned sunshade restores mean temperature and rainfall simultaneously in an idealized model.
- Emerging technologies could preferentially scatter sunlight in the near-infrared, providing a spectral sunshade.

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This article has been accepted for publication and undergone full peer review but has not been through the copyediting, typesetting, pagination and proofreading process, which may lead to differences between this version and the [Version of Record](#). Please cite this article as [doi: 10.1029/2020GL090876](https://doi.org/10.1029/2020GL090876).

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Abstract

Solar Radiation Modification (SRM) reduces the CO₂-induced change to the mean global hydrological cycle disproportionately more than it reduces the CO₂-induced increase in mean surface temperature. Thus if SRM were used to offset all CO₂-induced mean warming, global-mean precipitation would be less than in an unperturbed climate. Here we show that the mismatch between the mean hydrological effects of CO₂ and SRM may partly be alleviated by spectrally tuning the SRM intervention (reducing insolation at some wavelengths more than others). By concentrating solar dimming at near-infrared wavelengths, where H₂O has strong absorption bands, the direct effect of CO₂ on the tropospheric energy budget can be offset, which minimizes perturbations to the mean hydrological cycle. Idealized cloud-resolving simulations of radiative-convective equilibrium confirm that spectrally-tuned SRM can simultaneously maintain mean surface temperature and precipitation at their unperturbed values even as large quantities of CO₂ are added to the atmosphere.

Plain Language Summary

It may be possible to partly counteract CO₂-driven climate change by solar radiation modification (SRM) that intentionally reduces the amount of sunlight absorbed by the Earth. But different wavelengths of the solar spectrum are absorbed at different altitudes within the surface-atmosphere system, so different climatic effects are to be expected depending on which wavelengths of sunlight are affected by an SRM intervention. Here we show that if the goal is to minimize perturbations to the mean hydrological cycle, the ideal spectrally-tuned SRM intervention may need to focus on near-infrared wavelengths. This study clarifies the basic physics underlying the effects of SRM on atmospheric energetics and the mean hydrological cycle.

1 Introduction

Solar Radiation Modification (SRM) proposals aim to counteract CO₂-driven climate change by reducing the amount of sunlight absorbed by the Earth (D. Keith, 2013). Although significant scientific, practical, and ethical questions remain (e.g., P. J. Irvine et al., 2016; Preston, 2013), a growing body of evidence supports the notion that SRM could reduce many climatic changes that normally accompany a rise in CO₂ (e.g., Govindasamy & Caldeira, 2000; P. Irvine et al., 2019). Yet the interventions proposed for SRM

45 do not exactly counteract the radiative impacts of CO₂ forcing, so SRM would not ex-
46 actly offset CO₂-driven climate change. For example, a robust feature identified in sim-
47 ulations of geo-engineered climates is a weakened global hydrological cycle (Kravitz, Caldeira,
48 et al., 2013). That is, simulations of climates with high CO₂ and a dimmer sun have lower
49 mean precipitation and evaporation than do unperturbed climates with the same global-
50 mean temperature but lower CO₂ and a brighter sun (e.g., Bala et al., 2008; Tilmes et
51 al., 2013; Smyth et al., 2017).

52 The cause of the damped hydrological cycle in geo-engineered climates is well un-
53 derstood in terms of atmospheric energetics (e.g., Bala et al., 2008; Kravitz, Rasch, et
54 al., 2013; Kleidon et al., 2015). All else being equal, adding CO₂ to the atmosphere re-
55 duces the longwave (LW) cooling of the troposphere, and since the radiative cooling of
56 the troposphere is balanced primarily by latent heat released in precipitating clouds, a
57 reduction in radiative cooling leads to a reduction in precipitation (e.g. Allen & Ingram,
58 2002; Andrews et al., 2010). This is one of the “direct effects” of CO₂, so-called because
59 they are not mediated by changes in surface temperature (Dinh & Fueglistaler, 2017; Romps,
60 2020). Since the direct effect of CO₂ on tropospheric radiative cooling has remained largely
61 uncompensated in the SRM interventions that have been modeled so far, reductions in
62 mean precipitation have been identified as a robust feature of geo-engineered climates
63 (Kravitz, Caldeira, et al., 2013).

64 But, is this reduction in mean precipitation really an inevitable outcome of SRM
65 interventions? The purpose of this study is to explore, theoretically and in the context
66 of an idealized model, the possibility of a more complete radiative antidote to CO₂ forc-
67 ing — an antidote that simultaneously maintains mean temperature and precipitation
68 at their unperturbed values even as CO₂ is added to the atmosphere. Our approach ex-
69 ploits the fact that the shortwave (SW) opacity of the troposphere is not evenly distributed
70 across the solar spectrum, which means that different wavelengths of sunlight deposit
71 their energy within different layers of the coupled surface-troposphere system (Haigh et
72 al., 2010). This allows a spectrally-tuned SRM intervention (i.e., a wavelength-dependent
73 reduction in insolation) to restore energy balance at the tropopause and at the surface
74 simultaneously. As a result, spectrally-selective SRM could be less disruptive to the mean
75 hydrological cycle than spectrally-uniform SRM, which is the style of intervention that
76 has been modelled by the majority of previous studies (e.g., the “G1” experiment from
77 the recent GeoMIP project; Kravitz et al., 2011).

It is not currently known to what extent changes to regional precipitation in ge-engineered climates result from the physics we explore here — that is, from the mismatch between the effects of CO₂ and SRM on the tropospheric energy budget. Other factors, such as spatial or seasonal mismatches between CO₂ and SRM forcing (e.g., Lutsko et al., 2020), or stratospheric heating due to aerosol injection (Simpson et al., 2019), may dominate over the mean-energetics perspective that is the focus of our study. Future work investigating the effects of spectrally-tuned SRM interventions with a comprehensive global climate model will help to answer this important question. While our focus is on an idealized model, we review progress toward achieving spectrally-selective SRM in the real world, and note that the methods and physical insight gained from our approach are relevant to all SRM because no SRM intervention would be exactly spectrally uniform.

2 Theory

The use of SRM interventions to counteract CO₂ forcing is motivated by the standard forcing-feedback framework for temperature change (e.g., S. C. Sherwood et al., 2015), which states that the equilibrium change in surface temperature, ΔT_s , produced by an external perturbation is proportional to the radiative forcing at the tropopause produced by that perturbation, F_t (W/m²):

$$\Delta T_s = \alpha_T F_t, \quad (1)$$

where the constant of proportionality α_T (K/W/m²) is known as the “climate sensitivity” parameter. In this context, radiative forcing refers to the change in net radiative flux produced by the perturbation itself (i.e., before any adjustments in surface temperature). Motivated by equation (1), the SRM interventions previously modeled in the literature have aimed to lessen CO₂-induced warming by offsetting the (positive) CO₂ radiative forcing at the tropopause with a countervailing (negative) radiative forcing at the tropopause from SRM.

While the energy budget at the tropopause controls changes in surface temperature, it alone does not constrain the precipitation rate because changes in precipitation ΔP (kg/m²/s) are driven by changes in *tropospheric* radiative cooling, ΔQ (W/m²):

$$\Delta P = -\alpha_P \Delta Q, \quad (2)$$

105 where α_P (kg/J) is a “hydrological sensitivity” parameter¹ (e.g., O’Gorman et al., 2011;
 106 Pendergrass & Hartmann, 2014), and where negative values of Q indicate the typical sit-
 107 uation of net tropospheric radiative cooling. Since tropospheric radiative cooling depends
 108 on surface temperature as well as external perturbations such as increased CO₂ or changes
 109 in insolation, it is useful to separate ΔQ into the component produced by external per-
 110 turbations and the component that depends explicitly on ΔT_s (Lambert & Faull, 2007):

$$\Delta Q = F_a + \frac{\partial Q}{\partial T_s} \Delta T_s. \quad (3)$$

111 Here the external perturbation component F_a can be regarded as a radiative forcing of
 112 the troposphere, which is simply the difference between the radiative forcing of the ex-
 113 ternal perturbation evaluated at the tropopause and at the surface:

$$F_a = F_t - F_s. \quad (4)$$

114 From equations (1–4), we can deduce that a “radiative antidote” to CO₂ that maintains
 115 $\Delta T_s = \Delta P \simeq 0$ must offset the CO₂ radiative forcing at the tropopause and at the
 116 surface simultaneously.

117 Is such a radiative antidote to CO₂ possible? The left column of Figure 1 shows
 118 a schematic depiction of the radiative forcings produced by increased CO₂: a positive
 119 forcing at the tropopause $F_t^{\text{CO}_2}$, and a smaller positive forcing at the surface $F_s^{\text{CO}_2}$. For
 120 a given perturbation x , it is convenient to define a measure of how suppressed the as-
 121 sociated radiative forcing is at the surface compared to at the tropopause:

$$\epsilon_x \equiv F_s^x / F_t^x. \quad (5)$$

122 For example, for a CO₂ perturbation, the global-mean surface forcing is about half as
 123 large as the global-mean tropopause forcing (i.e., $\epsilon_{\text{CO}_2} \simeq 0.5$; Huang et al., 2017).

124 How do the surface and tropopause forcings compare for SRM interventions? The
 125 answer depends on 1) the shortwave opacity of the troposphere, and 2) the spectral sig-
 126 nature of the SRM intervention. The most commonly-studied SRM intervention is the
 127 idealized “sunshade” experiment (“G1”) from the GeoMIP protocol (Kravitz et al., 2011).
 128 This experiment calls for a simple reduction in the solar constant, which is implemented
 129 in numerical models as a spectrally-uniform fractional reduction in downwelling short-
 130 wave at the top-of-atmosphere (TOA). For this style of spectrally-flat SRM, ϵ_{SRM} de-

¹ Note that others have defined the hydrological sensitivity as $\Delta P / \Delta T_s$ (e.g., Kleidon et al., 2015).

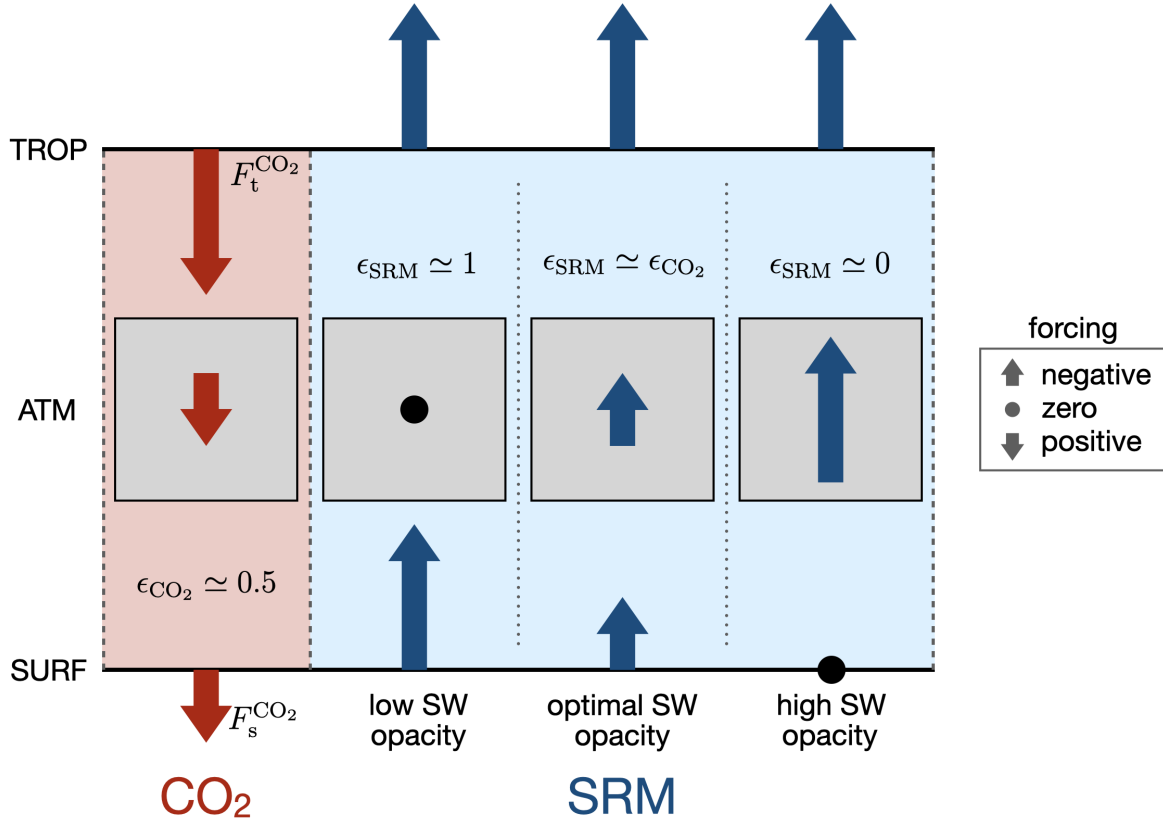


Figure 1. A schematic depiction of the radiative forcings at the tropopause (TROP), surface (SURF), and within the troposphere (ATM) produced by increasing CO₂ (leftmost column), or by idealized “sunshade” SRM interventions (three columns at right) in the style of the G1 experiment from GeoMIP (Kravitz et al., 2011). The SRM case is split into three sub-cases with bulk tropospheric shortwave opacities increasing from left to right. For a given perturbation x , ϵ_x is the ratio between the associated radiative forcing at the surface and at the tropopause: $\epsilon_x \equiv F_s^x / F_t^x$. Standard G1-style SRM falls toward the low end (left column) of the SW opacity scale.

131 depends on the bulk shortwave opacity of the troposphere: for a troposphere that is com-
 132 pletely (i.e., at all wavelengths) transparent to sunlight, $F_t^{\text{SRM}} = F_s^{\text{SRM}}$ and $\epsilon_{\text{SRM}} =$
 133 1, whereas for a troposphere that is completely opaque to sunlight, F_t^{SRM} is finite while
 134 $F_s^{\text{SRM}} = 0$, and so $\epsilon_{\text{SRM}} = 0$ (Figure 1, columns 2 and 4).

135 The above considerations allow us to understand what happens to the hydrolog-
 136 ical cycle when a CO₂ perturbation is combined with a sunshade-type SRM interven-
 137 tion. Combining equations (1–5), and setting $F_t^{\text{CO}_2} = -F_t^{\text{SRM}}$ as specified by the G1

138 GeoMIP protocol, we obtain the following expression for the change in precipitation:

$$\Delta P = \alpha_P F_t^{\text{CO}_2} (\epsilon_{\text{CO}_2} - \epsilon_{\text{SRM}}). \quad (6)$$

139 Hence the only circumstance in which spectrally-flat SRM can counteract CO₂ forcing
 140 at the tropopause and at the surface simultaneously is if the troposphere happens to have
 141 the correct intermediate bulk shortwave opacity so that $\epsilon_{\text{SRM}} = \epsilon_{\text{CO}_2}$. Otherwise, if a
 142 spectrally-flat SRM intervention is scaled such that F_t^{SRM} completely counteracts $F_t^{\text{CO}_2}$,
 143 there will be a residual radiative forcing of the troposphere. This residual forcing will
 144 drive a change in convective enthalpy fluxes from the surface (e.g., Dinh & Fueglistaler,
 145 2017), and perturb the hydrological cycle according to equations (2–3), as has been ob-
 146 served in the GeoMIP G1 experiment (Kravitz, Caldeira, et al., 2013; Tilmes et al., 2013).
 147 From the fact that the G1 experiment has yielded *reduced* mean precipitation (i.e., $\Delta Q >$
 148 0), we can deduce that the bulk shortwave opacity of Earth’s contemporary troposphere
 149 is too low (i.e., $\epsilon_{\text{SRM}} > \epsilon_{\text{CO}_2}$) for a spectrally-flat solar dimming to offset the direct ef-
 150 fect of CO₂ on tropospheric radiative cooling. By contrast, a spectrally-tuned SRM in-
 151 tervention could, in principle, concentrate the solar dimming in a portion of the solar
 152 spectrum with above-average tropospheric shortwave opacity, thereby allowing $\epsilon_{\text{SRM}} \simeq$
 153 ϵ_{CO_2} by construction. This is the basic insight underlying our suggestion that spectrally-
 154 tuned solar dimming could be a more complete radiative antidote to CO₂ forcing.

155 To quantitatively explore the potential of spectrally-tuned solar dimming, let us
 156 split the downwelling shortwave radiation at the tropopause, S_t^\downarrow , into N bands indexed
 157 by i , each with incoming power S_i (W/m²):

$$S_t^\downarrow = \sum_{i=1}^N S_i \quad (7)$$

158 In each of these bands, we denote the tropospheric transmissivity to vertically-propagating
 159 radiation as $\mathcal{T}_i = e^{-\tau_i}$, where τ_i is the total tropospheric column SW optical depth²
 160 in band i (assumed to be uniform within the band). If S_i is reduced by some fraction,
 161 the ratio of the associated forcings at the surface and tropopause is found, via Beer’s law,
 162 to be

$$\epsilon_i = \frac{\mathcal{T}_i^{1/\bar{\mu}}(1 - a_i)}{1 - a_i \mathcal{T}_i^{1/\bar{\mu} + D}}, \quad (8)$$

² For simplicity, here we assume SW attenuation is due only to molecular absorption, which is true for clear skies at wavelengths where Rayleigh scattering is negligible (e.g., the near-infrared).

where $\bar{\mu}$ is the effective cosine of the solar zenith angle, $D = 1.5$ is a two-stream hemispheric diffusivity factor (Clough et al., 1992), and where we have assumed a Lambertian surface with a band-specific SW albedo a_i . The optimal shortwave optical depth τ^* for offsetting CO₂ forcing is found by setting ϵ_i as given by equation (8) equal to ϵ_{CO_2} . To get a rough sense of the numbers, we can take $\epsilon_{\text{CO}_2} = 0.5$ and $a_i = 0$, yielding

$$\tau^* = \bar{\mu} \ln(2) \simeq 0.46, \quad (9)$$

where we have assumed $\bar{\mu} = 2/3$ (as is appropriate for the global mean; Cronin, 2014). For $a_i \neq 0$, τ^* is easily obtained via a rootsolver.

Now suppose we reduce the downwelling SW at the tropopause by band-specific fractional amounts γ_i , for $0 \leq \gamma_i \leq 1$ ($\gamma_i = 0$ corresponds to no reduction at the tropopause, whereas $\gamma_i = 1$ corresponds to complete blocking). The challenge of spectrally-tuned solar dimming amounts to finding a set of γ_i (i.e., a spectral filter) that simultaneously offsets $F_t^{\text{CO}_2}$ and $F_s^{\text{CO}_2}$, which is equivalent to simultaneously solving the following two equations:

$$F_t^{\text{CO}_2} = \sum_{i=1}^N \gamma_i S_i (1 - a_i \mathcal{T}_i^{1/\bar{\mu}+D}), \quad (10)$$

$$F_s^{\text{CO}_2} = \sum_{i=1}^N \gamma_i S_i \mathcal{T}_i^{1/\bar{\mu}} (1 - a_i). \quad (11)$$

To that end, it is instructive to consider a few limiting cases:

1. Filtering a band that passes through the atmosphere unabsorbed ($\mathcal{T}_i = 1$) perturbs the tropopause and surface energy budgets by the same amount, $\gamma_i S_i (1 - a)$.
2. Filtering a band that is completely absorbed in the troposphere ($\mathcal{T}_i = 0$) perturbs the tropopause energy budget by $\gamma_i S_i$, while leaving the surface energy budget unaffected.
3. Filtering a band for which $\mathcal{T}_i^{1/\bar{\mu}} (1 - a) / [1 - a \mathcal{T}_i^{1/\bar{\mu}+D}] = \epsilon_{\text{CO}_2}$ offsets the same *fraction* of CO₂ forcing at the tropopause and at the surface.

These principles suggest that there are many equally valid algorithms that could be used to design a spectral SRM filter. For simplicity, in this work we will simply find a contiguous band of wavenumbers that happens to have the correct distribution of optical depths to simultaneously solve (10–11).

189 3 Experimental methods

190 Our understanding of the effect of SRM on global-mean precipitation is based on
191 a radiative-convective equilibrium (RCE) perspective on the tropospheric energy bud-
192 get (Bala et al., 2008; Kravitz, Rasch, et al., 2013; Kleidon et al., 2015). The state of RCE
193 is the simplest system that faithfully captures the vertically-resolved energy budget of
194 Earth’s troposphere — that is, the balance between radiative cooling and convective heat-
195 ing. Therefore, the RCE framework is a natural testbed for a proof-of-principal demon-
196 stration of spectral SRM. We conducted RCE simulations with the cloud-resolving model
197 DAM (Romps, 2008), which has been used extensively to study tropical convection in
198 Earth’s atmosphere (e.g., Romps & Kuang, 2010; Romps, 2011, 2014; Seeley & Romps,
199 2015, 2016; Seeley, Jeevanjee, Langhans, & Romps, 2019; Seeley, Jeevanjee, & Romps,
200 2019). The default radiation scheme in DAM is RRTM (Clough et al., 2005; Iacono et
201 al., 2008), a correlated- k code, but for the purpose of this study we have coupled DAM
202 to a clear-sky radiation scheme based on the PCM-LBL code (Wordsworth et al., 2017).
203 This radiation scheme simply integrates the radiative transfer equation on a user-supplied
204 spectral grid using standard molecular opacity data from the HITRAN database (Gor-
205 don et al., 2017), a “brute-force” (i.e., wavenumber-by-wavenumber) approach to radi-
206 ation that greatly facilitates the investigation of spectrally-tuned solar dimming with-
207 out meaningful reductions in accuracy. We have benchmarked our radiation scheme against
208 RRTM and a line-by-line radiation code for an appropriate range of clear-sky conditions
209 and find very good agreement (Figures S2-3). In addition, whereas DAM typically uses
210 the Lin-Lord-Krueger bulk microphysics scheme (Lin et al., 1983; Lord et al., 1984; Krueger
211 et al., 1995), for this study we use the simplified cloud microphysics parameterization
212 described in Seeley, Jeevanjee, & Romps (2019). Since we adopt a clear-sky perspective
213 here, our results are not sensitive to the microphysics scheme, and we believe that the
214 simplified treatment of microphysics is appropriate for the present study, which is intended
215 simply as a proof-of-principle. Our simulations were conducted with a horizontal reso-
216 lution of $\Delta x = 2$ km on a square domain with a side length of approximately 100 km
217 and doubly-periodic horizontal boundary conditions. For further details regarding the
218 numerical modelling configuration, see the Supporting Information.

219 We first ran a control experiment (referred to as “CTRL”) with a total solar irra-
220 diance (TSI) of 510.375 W/m^2 and a fixed cosine of the solar zenith angle of $\bar{\mu} = 2/3$
221 (Cronin, 2014), yielding a downwelling shortwave flux at the TOA of 340.25 W/m^2 ; this

222 value matches the planetary-mean insolation $\mathcal{S}_0/4$, where $\mathcal{S}_0 = 1361 \text{ W/m}^2$ is the so-
 223 lar constant. The CTRL simulation was specified to have a preindustrial CO_2 concen-
 224 tration of 280 ppm and no ozone. CTRL was initialized from a similar RCE simulation
 225 over a fixed sea surface temperature and run for 1 year over a slab-ocean surface with
 226 a wavelength-independent albedo of 0.285, infinite horizontal conductivity (i.e., a uni-
 227 form temperature), and heat capacity equivalent to a layer of liquid water of depth 20
 228 cm. Results for CTRL were averaged over the final 200 days of model time. The equi-
 229 librated state of CTRL has a slab-ocean temperature of 288.64 K and mean precipita-
 230 tion rate of 3.17 mm/day. We then branched three experiments from the equilibrated
 231 state of CTRL: an abrupt quadrupling of CO_2 (referred to as “ $4\times\text{CO}_2$ ”), and two ex-
 232 periments for which the CO_2 quadrupling was accompanied by the application of some
 233 type of SRM. These branched simulations were run for an additional 3 years of model
 234 time, with results averaged over the final 100 days.

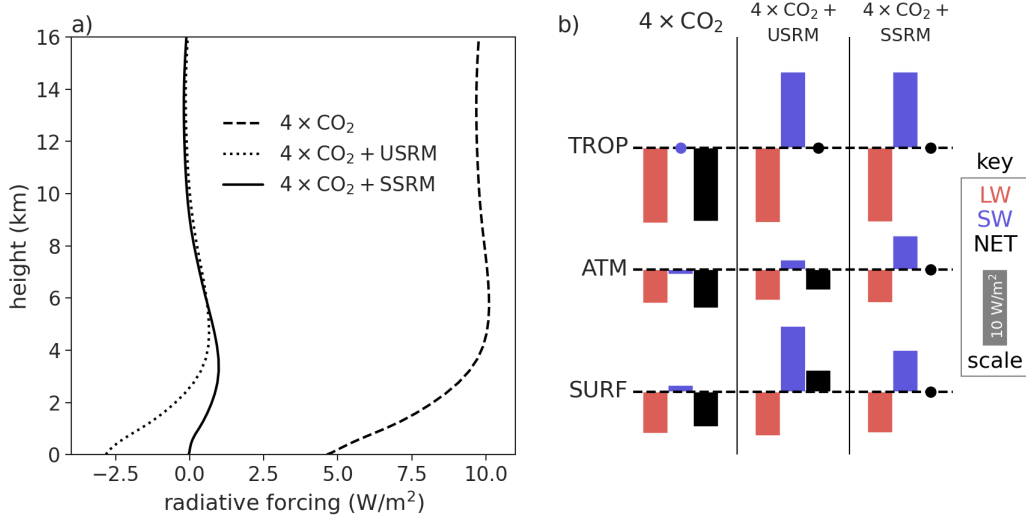


Figure 2. (a) Vertically-resolved radiative forcings, diagnosed as differences in net radiative fluxes. The forcings are shown for the three branched simulations with quadrupled CO_2 , two of which also include an SRM interventions (USRM or SSRM) as described in the main text. (b) Radiative forcings evaluated at the tropopause (top row, TROP; $z \simeq 15 \text{ km}$), surface (bottom row, SURF), and within the troposphere (middle row, ATM = TROP - SURF). For each simulation and level, the longwave (LW), shortwave (SW), and net (LW+SW) forcings are color-coded. By convention, positive forcings are depicted as downward-pointing bars, with the scale indicated by the gray 10 W/m^2 bar shown in the key; forcings with magnitude less than 0.2 W/m^2 are depicted as filled circles.

235 The SRM interventions were designed according to the principles discussed in sec-
236 tion 2. We first calculated the instantaneous radiative forcing from a quadrupling of CO₂
237 by double-calling the radiative transfer scheme at every radiative time step of the CTRL
238 simulation (once with 280 ppm CO₂ and once with 1120 ppm CO₂), and taking the dif-
239 ference between the net radiative fluxes. We evaluated these forcings at the tropopause
240 and at the surface; the tropopause was identified as the level at the top of the troposphere
241 where the time-averaged cloud fraction in CTRL falls below 1% (an altitude of approx-
242 imately 15 km). For the CO₂ quadrupling, the instantaneous radiative forcing at the tropopause
243 was found to be $F_t = 9.71 \text{ W/m}^2$, while the forcing at the surface was $F_s = 4.63 \text{ W/m}^2$
244 (Fig. 2; see also Table 1). Note that this implies $\epsilon_{\text{CO}_2} \simeq 0.5$ in our CTRL experiment,
245 close to the global-mean value reported in the literature (Huang et al., 2017). Strictly
246 speaking, the forcing F_t that enters into the forcing-feedback framework of equation (1)
247 should be the so-called *adjusted* forcing, which is the radiative flux imbalance at the tropopause
248 after stratospheric temperatures adjust to return the stratosphere to radiative equilib-
249 rium (e.g., Smith et al., 2018). For simplicity, here we use the instantaneous forcing in
250 place of the adjusted forcing, which was not found to be a large source of error.

251 In accordance with the GeoMIP G1 experiment (Kravitz et al., 2011), our first SRM
252 intervention was designed to completely offset the CO₂ radiative forcing at the tropopause
253 by reducing the solar constant. Since this amounts to a spectrally-uniform reduction in
254 TOA downwelling shortwave, we will refer to this intervention as USRM (with the “U”
255 indicating that this intervention is spectrally-uniform). The net shortwave flux at the
256 tropopause in CTRL is $S_t = 259.0 \text{ W/m}^2$, so we reduced the TSI by the factor $F_t/S_t =$
257 3.75% (from 510.375 W/m^2 to 491.23 W/m^2). To assess the efficacy of the intervention,
258 we averaged the radiative fluxes over the first week of the branched simulation, and cal-
259 culated radiative forcings as differences between these radiative fluxes and the mean ra-
260 diative fluxes from CTRL. As can be seen in the middle column of Figure 2b ($4\times\text{CO}_2+\text{USRM}$),
261 this spectrally-uniform SRM intervention restores the energy budget at the tropopause
262 (i.e., there is a negligible difference in net radiative flux at the tropopause between CTRL
263 and $4\times\text{CO}_2+\text{USRM}$), due to a cancellation between the positive LW forcing from the
264 CO₂ perturbation and the negative SW forcing of the SRM intervention. However, there
265 is a net forcing of the surface for this intervention, and equivalently, a change to the bulk
266 radiative flux divergence of the troposphere. The perturbation to the bulk tropospheric
267 radiative heating is positive (i.e., an anomalous heating) with magnitude $+2.72 \text{ W/m}^2$.

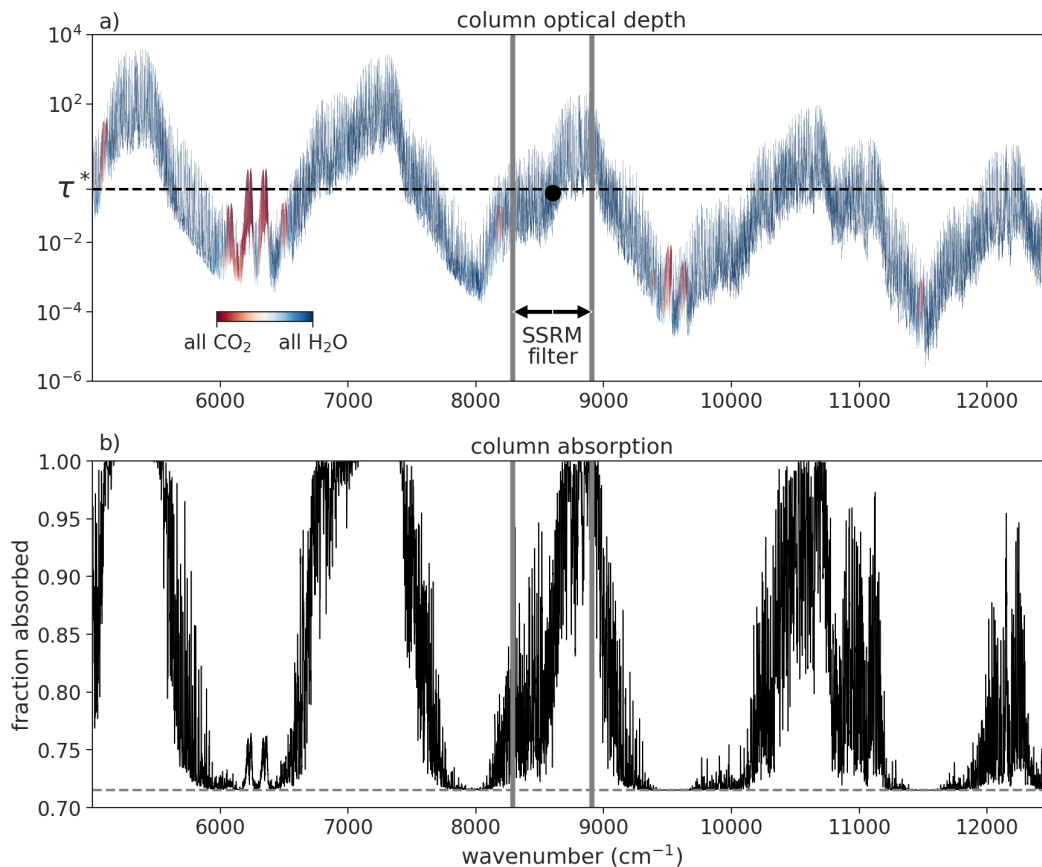


Figure 3. (a) Spectrally-resolved column optical depth from the CTRL experiment in the near-infrared. The data is color-coded according to the fraction of the surface optical depth contributed by H₂O versus CO₂. The optimal optical depth for offsetting CO₂ forcing at the tropopause and surface simultaneously, τ^* , is indicated by the horizontal dashed line. The spectral SRM filter spans the wavenumber range 8290–8910 cm^{-1} and is indicated by the gray bars. The (geometric) mean optical depth within the filtered band is indicated by the filled circle. (b) Spectrally-resolved column absorption from CTRL in the near-infrared. The surface co-albedo (i.e., 1 minus the surface albedo) is plotted as the horizontal dashed line, and sets the minimum column absorption (i.e., for a transparent atmosphere).

268 The LW effect of the CO₂ perturbation on the radiative cooling of the atmosphere is slightly
 269 larger than this, but is partially offset by a small anomalous shortwave cooling due to
 270 the USRM intervention.

271 The goal of spectrally-tuned solar dimming, on the other hand, is to *completely* off-
 272 set the direct effect of CO₂ by producing a larger anomalous shortwave cooling of the

273 troposphere. Here we suggest that this can be accomplished by concentrating the solar
 274 dimming in the near-infrared wavelengths (roughly 5000–12500 cm^{-1}), where H_2O has
 275 strong absorption bands that are primarily responsible for the shortwave heating of the
 276 troposphere. Specifically, to ensure a quantitatively accurate filter, we must choose γ_i
 277 to satisfy equations 10–11. By trial and error, we found that setting $\gamma_i = 1$ in the wavenum-
 278 ber range 8290–8910 cm^{-1} accomplishes this goal (Fig. 3). Note that the opacity in this
 279 band is attributable almost entirely to H_2O . There are other filters that also satisfy equa-
 280 tions 10–11, but we will take the filter shown in Figure 3 as our example of spectrally-
 281 tuned SRM (SSRM). The third column in Figure 2b shows that this filter, when com-
 282 bined with a quadrupling of CO_2 , produces no net forcing at the tropopause *or* surface,
 283 and therefore no anomalous bulk radiative heating of the troposphere. The filter works
 284 because it contains the correct balance of optically-thin and optically-thick wavelengths:
 285 although the optical depths within this band span roughly 4 orders of magnitude, the
 286 (geometric) mean optical depth within the filtered band is very close to the optimal op-
 287 tical depth $\tau^* = 0.34$ calculated by setting the right-hand side of equation (8) equal
 288 to ϵ_{CO_2} and solving for τ_i (Fig. 3a). The most optically-thick wavelengths within the fil-
 289 tered band are almost entirely absorbed within the troposphere, whereas the most optically-
 290 thin wavelengths are absorbed only at the surface (Fig. 3b).

291 Although the SSRM filter nullifies the direct effect of CO_2 on *bulk* tropospheric ra-
 292 diative heating, the vertically-resolved compensation is not exact (Fig. 3a). This slight
 293 redistribution of radiative heating rates in the vertical could, in principle, affect atmo-
 294 spheric dynamics.

295 4 Results

296 We have seen in Figures 2-3 that it is possible to design a spectrally-tuned SRM
 297 intervention that offsets the radiative forcing from CO_2 at the tropopause and at the sur-
 298 face simultaneously. But, does this SSRM approach outperform the USRM approach at
 299 the task of maintaining mean temperature and precipitation at their unperturbed val-
 300 ues? Figure 4 shows time series of surface temperature and precipitation from the branched
 301 RCE experiments. The surface warms rapidly in the $4\times\text{CO}_2$ experiment, eventually equi-
 302 librating at a surface temperature warmer by $\Delta T_s = 6.89$ K after approximately 3 years
 303 of model time. Therefore, the equilibrium climate sensitivity for our model configura-
 304 tion is approximately 3.5 K, squarely within the best-estimate range for ECS (S. Sher-

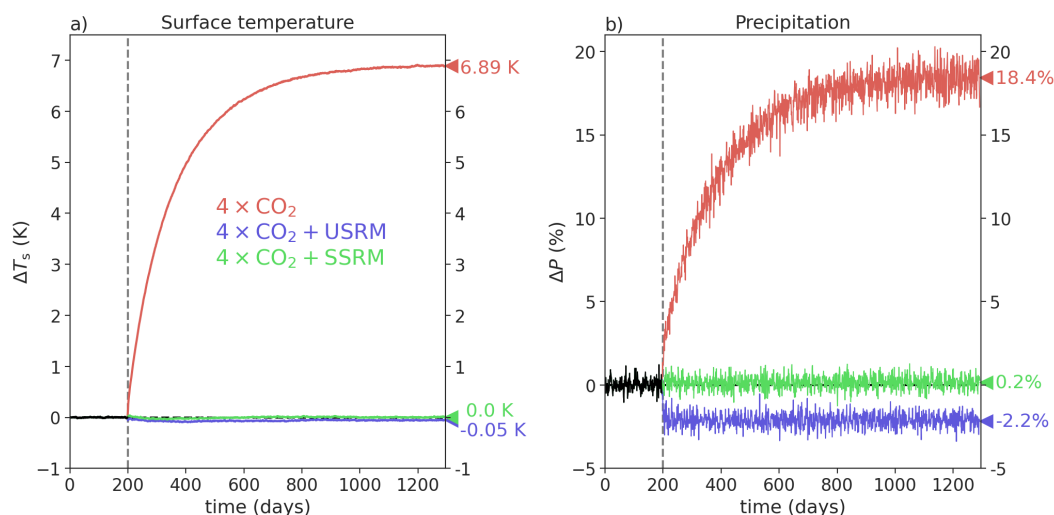


Figure 4. Anomaly timeseries of (a) slab-ocean temperature T_s and (b) precipitation rate P from the DAM experiments. The three experiments with quadrupled CO_2 are branched from CTRL at day 200 and run for 3 additional years of model time. The precipitation timeseries is plotted as a moving average with a centered window of 1 week to reduce noise. Quantities averaged over the final 100 days of the simulations are marked on the ordinates at right.

wood et al., 2020). This large warming causes an increase in mean precipitation of 18.4%,
 or roughly 2.7 %/K, which is the expected effect of a deepening troposphere under warm-
 ing (Jeevanjee & Roms, 2018).

Both SRM interventions (USRM and SSRM) greatly reduce the magnitude of changes
 to temperature and precipitation. For surface temperature, the two interventions are roughly
 equally effective: both limit the magnitude of changes in surface temperature to ≤ 0.05
 K, more than two orders of magnitude smaller than the warming caused by quadrupling
 CO_2 without any form of SRM. For precipitation, however, the effectiveness of the SRM
 interventions differs greatly: USRM causes a decrease in mean precipitation of -2.2%,
 whereas SSRM limits the change in precipitation to less than +0.2%. Therefore, the SSRM
 intervention is indeed a more complete radiative antidote to CO_2 forcing than the USRM
 intervention, because it nullifies both the greenhouse effect of CO_2 on surface temper-
 ature and the direct effect of CO_2 on precipitation. These results are summarized in Ta-
 ble 1.

	4×CO ₂	4×CO ₂ +USRM	4×CO ₂ +SSRM
F_t (W/m ²)	9.71	-0.105	-0.155
F_s (W/m ²)	4.63	-2.82	-0.04
ΔT_s (K)	6.89	-0.055	0.005
ΔP (%)	18.4	-2.16	0.157

Table 1. Radiative forcings at the tropopause (F_t) and surface (F_s), as well as mean changes in surface temperature T_s and precipitation P , from the three DAM experiments with quadrupled CO₂.

5 Discussion

In this study, we have developed the theory of spectrally-tuned SRM interventions. Such interventions have the goal of simultaneously maintaining mean surface temperature and mean precipitation at their unperturbed values even as large quantities of CO₂ are added to the atmosphere. Theoretically, this is made possible by the strong absorption bands of H₂O in the near-infrared: by concentrating solar dimming at these wavelengths, it is possible to produce an anomalous shortwave cooling of the troposphere that offsets the longwave heating of additional CO₂. Equivalently, a successful spectrally-tuned solar dimming preserves the energy budget of the tropopause and the surface (equations 10–11), whereas spectrally-flat solar dimming can preserve the energy budget at the tropopause but leaves the surface energy budget perturbed.

As a proof-of-principle, we have demonstrated the success of spectrally-tuned SRM in idealized cloud-resolving model experiments. Although we have only investigated SSRM in a configuration that entirely offsets CO₂ forcing at the tropopause (in accordance with the GeoMIP G1 protocol; Kravitz et al., 2011), our results can be generalized to help understand the effects of SRM interventions that offset only a fraction of CO₂ forcing. Suppose that an SRM intervention is designed such that $F_t^{\text{SRM}} = -\beta F_t^{\text{CO}_2}$, for $0 \leq \beta \leq 1$ (i.e., $\beta = 0$ corresponds to no offsetting of CO₂ forcing, while $\beta = 1$ corresponds to complete offsetting). Combining equations (1–5), we obtain the following expression for the change in precipitation, which generalizes equation (6):

$$\Delta P = -\alpha_P F_t^{\text{CO}_2} \left[\underbrace{(1 - \beta) + \beta \epsilon_{\text{SRM}} - \epsilon_{\text{CO}_2}}_{\text{direct effect}} + \underbrace{\frac{\partial Q}{\partial T_s} \alpha_T (1 - \beta)}_{\text{warming effect}} \right], \quad (12)$$

339 where we have identified with underbraces the two sources of changes in precipitation:
 340 1) the direct effect from the combination of a CO₂ perturbation and an SRM interven-
 341 tion, and 2) the effect of warming on precipitation. By putting in representative num-
 342 bers, we can use equation (12) to make several useful observations. Consider first CO₂
 343 forcing alone ($\beta = 0$). Taking $\epsilon_{\text{CO}_2} = 0.5$, $\frac{\partial Q}{\partial T_s} = -3 \text{ W/m}^2/\text{K}$ (Jeevanjee & Romps,
 344 2018), and $\alpha_T = 0.7 \text{ K/W/m}^2$ (as inferred from our $4 \times \text{CO}_2$ experiment, in which a \simeq
 345 10 W/m^2 tropopause forcing causes a $\simeq 7 \text{ K}$ warming), equation (12) suggests that the
 346 direct effect of CO₂ on precipitation is smaller than the warming effect by a factor of about
 347 $1/4$, close to the estimate of Romps (2020). This implies that, if the goal is to minimize
 348 disruption to the hydrological cycle, minimizing changes in surface temperature via the
 349 tropopause energy budget is the most powerful lever. It is only when $\beta \simeq 1$, as in our
 350 experiments and the G1 experiment protocol, that the warming effect is suppressed enough
 351 to allow the direct effect to dominate changes in precipitation; in this limit, equation (12)
 352 shows that the direct effect is controlled by the difference $\epsilon_{\text{SRM}} - \epsilon_{\text{CO}_2}$, as previously
 353 discussed in section 2 (c.f. equation 6, Fig. 1). In this limit, if the goal is to minimize
 354 the perturbation to the mean precipitation rate, the spectral properties of the SRM in-
 355 tervention should be tuned so that $\epsilon_{\text{SRM}} = \epsilon_{\text{CO}_2}$. In general, the optimal value of ϵ_{SRM}
 356 will depend on the strength of the SRM intervention as well as the objective function.

357 Given the success of spectral SRM in our idealized model, it is natural to wonder
 358 how spectral SRM might be realized in the real world. At present, there is no off-the-
 359 shelf commercial technology that could be used to implement spectral SRM without pro-
 360 hibitive costs and environmental impacts. Yet, SRM would likely be implemented over
 361 a time scale of a century or more, so there is time for technological innovation, and al-
 362 ready there are signs that “designer materials” with tuneable extinction coefficients at
 363 near-infrared wavelengths may be within reach. Metallic nanoparticles that exhibit op-
 364 tical plasmonic resonance (Khlebtsov & Dykman, 2010) can exhibit narrow-band scat-
 365 tering or absorption in the optical and near-infrared, with resonant spectral widths of
 366 order 1000 wavenumbers (Berkovitch et al., 2010) — consistent with the size of filter we
 367 analyze in this work. Diffractive structures and resonant scatterers for SRM were pro-
 368 posed over two decades ago (Teller et al., 1997); similarly, self-levitating atmospheric scat-

369 terers for SRM were proposed a decade ago (D. W. Keith, 2010), and are now being phys-
370 ically demonstrated in the lab (Cortes et al., 2020). A small but growing body of liter-
371 ature has explored space-based SRM since 1989, and several of these proposals exploit
372 diffractive screens (Angel, 2006). All of these methods could serve as the basis for spectrally-
373 tunable SRM interventions.

374 Even if it turns out that spectrally-tuned SRM technologies will never be practi-
375 cal or cheap enough for use, our results remain relevant to more mainstream approaches
376 to SRM (e.g., with stratospheric sulfate aerosols), for the simple reason that any real-
377 world implementation of SRM will not be spectrally uniform. We have shown how to map
378 the spectral characteristics of candidate SRM technologies onto their expected impacts
379 on mean precipitation, thereby providing a new metric for evaluating such technologies.
380 Indeed, prior work has shown that different SRM technologies have different effects on
381 precipitation rates (Niemeier et al., 2013). Our results indicate that one potential source
382 of such differences is the different spectral characteristics of the proposed technologies;
383 future work could assess this possibility with the aid of equations (10–11), which pro-
384 vide a quick method of parsing the “design space” of SRM technologies without resort-
385 ing to computationally-expensive simulations with global climate models.

386 Overall, although our results regarding the potential of spectral SRM are promis-
387 ing, many questions remain. It is important to realize that designing a radiative anti-
388 dote to CO₂ is substantially easier for atmospheres that are statistically homogeneous
389 in the horizontal (e.g., our RCE simulations). On the real Earth, spatial heterogeneity
390 in surface temperature, water vapor content, and albedo would cause the ideal spectral
391 SRM intervention itself to be spatially heterogeneous. Another weakness of the theory
392 of spectral SRM developed here is that the surface Bowen ratio is unconstrained, which
393 means that the precipitation rate could change even when the radiative energy budget
394 of the troposphere is unperturbed. This effect could be especially important in models
395 with heterogeneous surface conditions. Global models are the only way to assess changes
396 to regional precipitation, which are more relevant to society than the global-mean change.
397 For these reasons and more, future work should test the effectiveness of spectral SRM
398 in comprehensive global models. It seems possible that SRM interventions that are less
399 disruptive of the tropospheric energy balance will be less disruptive of the climate on a
400 regional scale, but further work is needed to verify this hypothesis.

Acknowledgments

Simulation data and Python source code for reproducing the figures in this manuscript is available at <https://doi.org/10.5281/zenodo.4035201>.

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